The Formation and Evolution of Exponential Disks in Galaxies

October 5-9, 2014
Lowell Observatory
Flagstaff, AZ, USA

LOC

Deidre Hunter

SOC

- Joss Bland-Hawthorn
- Bruce Elmegreen (co-chair)
- → Peter Erwin
- → Kambiz Fathi
- Annette Ferguson
- → Ken Freeman
- Philip Hopkins
- Roelof de Jong
- Eija Laurikainen (co-chair)
- Rok Roskar
- Heikki Salo
- → Patricia Tissera

Confirmed Speakers

- Roberto Abraham
- Lia Athanassoula
- Judit Bakos
- Joss Bland-Hawthorn
- Frédéric Bournaud
- → Chris Brook
- Giovanni Carraro
- → Simon Driver
- **▶ Peter Erwin**
- Kambiz Fathi
- → Sofia Feltzing
- Annette Ferguson
- → Ken Freeman
- Carme Gallart

- Ortwin Gerhard
- Roelof de Jong
 - Dusan Keres
 - Bärbel Koribalski
 - Elizabeth McGrath
 - Rok Roskar
 - Heikki Salo
 - > Sabastian Sanchez
 - Linda Tacconi
 - David Thilker
 - → Patricia Tissera
 - → Piet van der Kruit
 - → Stijn Wuyts
 - Rosemary Wyse
 - Peter Yoachim

Exponential stellar disks are ubiquitous. The stars in both spiral and dwarf galaxies are generally found to be organized in exponential disks, even to very low surface densities and in both stellar dominated and gas dominated galactic environments. But why is this? The associated gas disks do not fall off with radius in the same manner. Furthermore, star formation is highly lumpy. How does lumpy star formation produce distributions of stars that fall off smoothly? And how are these profiles maintained over many Gyr? In addition, abrupt breaks in the stellar surface brightness or density profiles are also common. So what happens at the break in these galaxies?

<u>Important Deadlines:</u>

until May 15: Pre-registration and suggestion of talks
May 15 - August 1: Registration and abstract submission
September 5: Cancellation with full refund
September 5: Guaranteed block hotel reservation



