ERRATUM: "HOW TO CONSTRAIN YOUR M DWARF: MEASURING EFFECTIVE TEMPERATURE BOLOMETRIC LUMINOSITY, MASS, AND RADIUS" (ApJ, 804, 64)

ANDREW W. MANN^{1,2,8,9}, GREGORY A. FEIDEN³, ERIC GAIDOS^{4,5,10}, TABETHA BOYAJIAN⁶, AND KASPAR VON BRAUN⁷

University of Texas at Austin, USA

² Institute for Astrophysical Research, Boston University, USA ³ Department of Physics and Astronomy, Uppsala University, Box 516, SE-751 20, Uppsala, Sweden ⁴ Department of Geology and Geophysics, University of Hawaii at Manoa, Honolulu, HI 96822, USA Max Planck Institut für Astronomie, Heidelberg, Germany ⁶ Department of Astronomy, Yale University, New Haven, CT 06511, USA Lowell Observatory, 1400 W. Mars Hill Rd., Flagstaff, AZ, USA

Received 2016 January 15; accepted 2016 January 15; published 2016 March 1

Due to a press error, in the original article, Tables 1, 2, and 3 were published with incorrect values. IOP sincerely regrets the error and has reproduced the tables in full below.

	Table 1 Mass and Radius Relations										
Y	Х	Eqn #	а	b	с	d	e	f	σ^{a} %	χ^2_{ν}	
R_*	M_{K_S}	(4)	1.9515	-0.3520	0.01680				2.89	0.93	
R_*	M_{K_S} , [Fe/H]	(5)	1.9305	-0.3466	0.01647			0.04458	2.70	0.88	
R_*	$T_{\rm eff}/3500$	(4)	10.5440	-33.7546	35.1909	-11.5928			13.4	2.35	
R_*	$T_{\rm eff}/3500$, [Fe/H]	(5)	16.7700	-54.3210	57.6627	-19.6994		0.4565	9.3	1.10	
M_{\star}^{c}	M_{K_S}	(10)	0.5858	0.3872	-0.1217	0.0106	-2.7262×10^{-4}		1.8	0.37	

Notes. For the first, third, and fifth equation $Y = a + bX + c^2$..., for the equations including [Fe/H] the right-hand side is multiplied by (1 + f[Fe/H]).

Notes. For the first, third, and nith equation I = a + vA + c ..., for the standard deviation of $\frac{R_{*,observed} - R_{*,predicted}}{R_{*,observed}}$. The last relation is quoted as the percent scatter in M_{*} . ^c Semi-empirical relation derived using empirical K_S-band magnitudes and masses estimated from our model analysis. Coefficients are calculated using maximum likelihood and a MCMC method. See Section 8 for details.

$T_{\rm eff}$ Relation Coefficients										
Y	Х	a	b	с	d	e	f	g	σ^{a} K	χ^2_{ν}
$T_{\rm eff}/3500$	BP - RP	3.245	-2.4309	1.043	-0.2127	0.01649			52	0.88
$T_{\rm eff} / 3500$	V - J	2.840	-1.3453	0.3906	-0.0546	0.002913			55	0.93
$T_{\rm eff} / 3500$	V - Ic	2.455	-1.5701	0.6891	-0.1500	0.01254			53	0.94
$T_{\rm eff} / 3500$	r-z	1.547	-0.7053	0.3656	-0.1008	0.01046			58	1.06
$T_{\rm eff}/3500$	r - J	2.445	-1.2578	0.4340	-0.0720	0.004502			58	1.04
$T_{\rm eff}/3500$	BP - RP, [Fe/H]	2.835	-1.893	0.7860	-0.1594	0.01243	0.04417		45	0.60
$T_{\rm eff} / 3500$	V - J, [Fe/H]	2.515	-1.054	0.2965	-0.04150	0.002245	0.05262		42	0.53
$T_{\rm eff} / 3500$	V - Ic, [Fe/H]	1.901	-0.6564	0.1471	-0.01274		0.04697		48	0.67
$T_{\rm eff}/3500$	r - z, [Fe/H]	1.572	-0.7220	0.3560	-0.09221	0.009071	0.05220		50	0.71
$T_{\rm eff}/3500$	r - J, [Fe/H]	2.532	-1.319	0.4449	-0.07151	0.004333	0.05629		47	0.63
$T_{\rm eff}/3500$	BP - RP, J - H	3.172	-2.475	1.082	-0.2231	0.01738	0.08776	0.04355	49	0.78
$T_{\rm eff} / 3500$	V - J, J - H	2.769	-1.421	0.4284	-0.06133	0.003310	0.1333	0.05416	48	0.71
$T_{\rm eff} / 3500$	V - Ic, J - H	1.568	-0.4381	0.07749	-0.005610		0.2441	-0.09257	52	0.85
$T_{\rm eff}/3500$	r-z, J-H	1.384	-0.6132	0.3110	-0.08574	0.008895	0.1865	-0.02039	55	0.90
$T_{\rm eff}/3500$	r-J, J-H	2.151	-1.092	0.3767	-0.06292	0.003950	0.1697	0.03106	52	0.79

Table 2

Note. The first five formulae follow Equation (4), the middle five follow Equation (6) (f is the coefficient of the [Fe/H] term), and the last five follow Equation (7) (f and g are the coefficients for the J - H and $(J - H)^2$ terms, respectively). Equations using J - H as an additional variable are meant for when the metallicity is not known. ^a We report the scatter in the predicted — observed (from spectrum) $T_{\rm eff}$. Conservatively, these errors should be added (in quadrature) with our typical spectroscopic uncertainty (60 K).

⁸ Harlan J. Smith Fellow.

⁹ Visiting Researcher.

¹⁰ Visiting Scientist.

BC _Y	Х	a	b	с	d	e	σ	χ^2_{ν}
V	V - J	0.5817	-0.4168	-0.08165	4.084×10^{-3}	•••	0.016	0.88
Rc	V - J	2.127	-1.059	0.1029	-7.881×10^{-3}		0.031	2.97
Ic	V-J	0.4440	0.2331	-0.05313			0.037	2.47
r	r - J	0.8958	-0.5081	-0.07387	3.999×10^{-3}		0.016	0.56
i	r - J	0.4431	-0.06470	-0.04038	2.798×10^{-5}		0.031	2.86
z	r-J	0.05373	0.2980	-0.05001		•••	0.035	3.53
Gaia	BP - RP	0.7384	-0.7398	0.01340			0.045	5.93
J	V - J	0.8694	0.3667	-0.02920			0.016	0.90
J	r - J	0.8790	0.5068	-0.07791	4.338×10^{-3}		0.016	0.92
Η	V - J	1.834	0.2054	-0.01271			0.030	1.96
Η	r - J	1.939	0.1969	-0.01337		•••	0.029	1.87
Κ	V - J	1.421	0.6084	-0.09655	6.263×10^{-3}	•••	0.036	2.44
Κ	r - J	1.719	0.5236	-0.09085	6.735×10^{-3}	•••	0.036	2.36
V	V - J, [Fe/H]	0.6570	-0.4710	-0.06943	3.206×10^{-3}	-0.04885	0.012	0.50
Rc	V - J, [Fe/H]	2.183	-1.102	0.1126	-8.579×10^{-3}	-0.09587	0.025	1.92
Ic	V - J, [Fe/H]	0.5043	0.1994	-0.04883		-0.06312	0.032	1.82
r	r - J, [Fe/H]	0.9341	-0.5432	-0.06423	3.170×10^{-3}	-0.05569	0.012	0.28
i	r - J, [Fe/H]	0.5235	-0.1326	-0.02203	-1.541×10^{-3}	-0.1396	0.028	2.60
z	r - J, [Fe/H]	0.1009	0.2658	-0.04509		-0.07352	0.028	2.44
Gaia	BP - RP, [Fe/H]	0.7567	-0.7541	0.01574		-0.1212	0.037	4.39
\overline{J}	V - J, [Fe/H]	0.8879	0.3563	-0.02791		-0.04857	0.012	0.64
J	r - J, [Fe/H]	0.9672	0.4291	-0.05677	2.528×10^{-3}	-0.05249	0.012	0.56
Η	V - J, [Fe/H]	1.796	0.2260	-0.01525		0.09544	0.021	1.02
Η	r - J, [Fe/H]	1.915	0.2135	-0.01582		0.09088	0.021	1.01
Κ	V - J, [Fe/H]	1.197	0.7714	-0.1339	8.998×10^{-3}	0.09572	0.030	1.68
Κ	r - J, [Fe/H]	1.572	0.6529	-0.1260	9.746×10^{-3}	0.08987	0.030	1.68

 Table 3

 Bolometric Correction Formulae

Note. All relations are of the form $BC_Y = a + bX + cX^2 + dX^3 + e([Fe/H])$, where Y is a the filter listed above, and X is the specified color.